# ON THE INTEGRABILITY OF THE HAMILTON - JACOBI EQUATION IN GENERALIZED COORDINATES***) 

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Invariant criteria of existence of the separable genexalized coordinates are obtained for certain classes of conservative systems, and examples are given.

By definition, the variables $q^{0}=t, q^{1}, \ldots, q^{n}$ in the Hamilton-Jacobi equation

$$
\begin{equation*}
p_{0}+H\left(q^{0}, q^{1}, \ldots, q^{n}, p_{1}, \ldots, p_{n}\right)=0, \quad\left(p_{0}=\frac{\partial V}{\partial t}, p_{1}=\frac{\partial V}{\partial q^{i}} ; i=1, \ldots, n\right) \tag{0.1}
\end{equation*}
$$

are (locally) separable if the equation admits a complete integral of the form

$$
\begin{equation*}
V=\sum_{i=0}^{n} V_{i}\left(q^{i}, c_{1}, \ldots, c_{n}\right)+c_{n+1} \tag{0.2}
\end{equation*}
$$

where $c_{1}, \ldots, c_{n}, c_{n+1}$ are the integration constants. The variables can be separated if and only if $/ 1,2 /$ the Hamiltonian function $H(t, q, p)$ satisfies identically the conditions

The above conditions are noninvariant with respect to the choice of canonical variables: the function $H$ can convert ( 0.3 ) into identities for one set of variables, and not satisfy (0.3) for another set. For example, in the Hamilton-Jacobi equation for the problem of two bodies the spherical coordinates are separable, while the Cartesian coordinates are not. A large body of literature exists dealing with the problem of integrability of equation (0.1) using the variable separation method (see detailed bibliography in $/ 3,4 /$ ). Interpreting the conditions (0.3) as differential equations in $H$, one can say that the majority of these investigations were limited to indicating some particular solutions of equations (0.3). Since the Hamiltonian function is known in advance for every concrete problem, another problem of practical interest arises, namely, that if the variables in (O.1) are not separable, then does a canonical transformation $(t, f, q, p) \rightarrow(\tau, K, Q, P)$ exist such that the variables $Q^{0}=\tau, Q^{1}, \ldots$, $Q^{n}$ in the transformed Hamilton-Jacobi equation $p_{0}+K(\tau, Q, P)=0$ are separable. In the general formulation the problem remains unsolved. The few constructive results refer to specific problems of dynamics (e.g. the problem of motion of a material point in a conservative force field /5-8/). Sufficient conditions are also obtained for which the integral $H=$ const is a unique first integral of the Hamilton's equations analytic over the whole configurational manifold of the system /9.10/ (from which it follows that the holonomic systems satisfying these conditions have globally no canonical transformation, referred to above, in the class of analytic functions). Let

$$
\begin{equation*}
H=\frac{1}{2} \sum_{i, j=1}^{n} a^{i j}(t, q) p_{i} p_{j}+\sum_{i=1}^{n} a^{i}(t, q) p_{i}+a(t, q) \tag{0.4}
\end{equation*}
$$

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The formulation of the problem in question can be made more specific in case of the systems with Hamiltonian function of the form (0.4). Namely, the solution of the following problem is of interest. How to use the form of the function ( 0.4 ) referred to the arbitrary generalized coordinates to find out whether separable coordinates

$$
Q^{i}=Q^{i}\left(t, q^{1}, \ldots, q^{n}\right)(i=1, \ldots, n)
$$

exist for this function. This formulation is also found insufficient for solving the problem. The author of /8/ found in this connection the most general relationships connecting the coefficients $a^{i j}, a^{i}$, and $a$ with the separable variables $t, q^{1}, \ldots, q^{n}$. The relations represent rational expressions in terms of a number of arbitrary functions of a single argument. In addition it was shown in $/ 8,11 /$ that in the case when separable generalized coordinates exist, special separable coordinates $x_{1}, \ldots, x_{n}$ (called in /ll/ "normal") can always be found in which the Hamiltonian function of the system assumes a particularly simple form

$$
\begin{equation*}
H^{*}-\frac{1}{2} \sum_{k, l=1}^{r} b^{k l}\left(t, x^{r+1}, \ldots, x^{n}\right) \pi_{k} \pi_{l}+\frac{1}{2} \sum_{v==r+1}^{n} b^{v x}\left(t, x^{r+1}, \ldots, x^{n}\right)\left(\pi_{v}\right)^{2}+\sum_{k=1}^{r} h^{k}\left(t, x^{r+1}, \ldots, x^{n}\right) \pi_{k}+b\left(t, x^{r+1}, \ldots, x^{n}\right) \tag{0.5}
\end{equation*}
$$

where $\pi_{j}$ are the impulses conjugate with respect to the normal coordinates. The number $0 \leqslant r \leqslant n$ characterizes the type of the separable system, and for $r=0$ the corresponding sums in the formula (0.5) vanish.

1. Let us consider a natural system with two degrees of freedom. Let $L=T+U$ be the Lagrange's function and $V_{2}$ the configurational Riemannian manifold of the system,

$$
\begin{equation*}
d s^{2}=2 T d t^{2}=\sum_{i, j=1}^{2} a_{i j} d q^{i} d q^{j} \tag{1.1}
\end{equation*}
$$

its metric, and let the components $a_{i j}$ have continuous partial derivatives up to and including the fifth order. We assume that the curvature $K$ of the manifold is nonconstant and satisfies the conditions

$$
\begin{gather*}
\Delta_{1} K=f(K), \quad \Delta_{2} K=F(K), \quad \Delta_{1} K=\sum_{i, j=1}^{2} a^{i j} \partial_{i} K \partial_{j} K, \quad \Delta_{2} K=\frac{1}{\delta} \sum_{i, j=1}^{2} \partial_{i}\left(\delta a^{i j} \partial_{j} K\right)  \tag{1.2}\\
\left\|a^{i j}\right\|=\left\|a_{i j}\right\|^{-1}, \quad \delta^{2}=a_{11} a_{22}-\left(a_{12}\right)^{2}, \quad \partial_{i}=\frac{\partial}{\partial_{q} q^{i}}
\end{gather*}
$$

The conditions (1.2) are necessary and sufficient /12/ for the equations of the geodesics in $V_{2}$ to have exactly one linear integral $J=$ const. Let us put $\omega=\exp \left(-\int F(K) / f(K) d K\right)$ and introduce the following new coordinates:

$$
\begin{equation*}
Q^{1}=K, \quad Q^{2}=\int \frac{\omega}{\delta}\left[\left(a_{21} \partial_{1} K-a_{11} \partial_{2} K\right) d q^{1}+\left(a_{22} \partial_{1} K-a_{12} \partial_{2} K\right) d q^{2}\right] \tag{1.3}
\end{equation*}
$$

in which the metric (1.1) assumes the form

$$
\begin{equation*}
\left.d s^{2}=\frac{1}{\omega^{2} f} I\left(\omega d Q^{1}\right)^{2}+\left(d Q^{2}\right)^{2}\right] \tag{1.4}
\end{equation*}
$$

The formulas (1.3) and (1.4) are defined in the neighborhood of an arbitrary regular point of the function $K$. We shall require that at least one of the following four relations (a prime denotes differentiation with respect to $K$ ) does not hold for the functions $f$ and $F$ :

$$
\begin{align*}
& F^{\prime} f+5 K f-f^{\prime \prime} f+\frac{1}{5}\left(f^{\prime}-F\right)\left(\frac{7}{2} f^{\prime}-F\right)=0  \tag{1.5}\\
& 3\left(4 F^{\prime}-\frac{3}{2} f^{\prime}\right)\left[5 K f+\left(F-\frac{1}{2} f^{\prime}\right)\left(F-f^{\prime}\right)\right]-25 K f^{2}=0 \\
& \frac{S^{\prime}}{5}+\frac{f^{\prime}}{f}-\frac{1}{K}-\frac{P}{S K}=\frac{P^{\prime}}{P}+\frac{f^{\prime}}{f}+\frac{R K}{f P}=\frac{R^{\prime}}{R}+\frac{f^{\prime}}{2!}-\frac{2(S+2 P)}{R} \\
& \left(P=\frac{1}{2} f^{\prime}-F, R=\frac{\left(6 F-7 / 2 f^{\prime}\right) f}{f^{\prime}-F}, \quad S=\frac{\left(F-1 / 2 f^{\prime}\right)\left(f^{\prime}-F\right)-5 K i}{f^{\prime}-F}\right)
\end{align*}
$$

Under this condition the equations of the geodesics in $V_{2}$ admit exactly two linearly independent quadratic integrals $/ 12,13 /$, i.e. any quadratic integral $I=$ const of the geodesicequations represents the combination

$$
\begin{equation*}
\left.I \equiv c_{1}\left(\frac{d s}{d t}\right)^{2} \right\rvert\, c_{4} J^{2} \quad\left(c_{1}, c_{2} \quad \text { are constants }\right) \tag{1.6}
\end{equation*}
$$

Theorem 1. Let the curvature $K$ of the manifold with the metric (1.1) be nonconstant and satisfy the conditions (1.2), and let at least one of the relations (1.5) be violated. Then the necessary and sufficient condition for the existence of separable generalized coordinates of the system near any regular point of the function $K$ is, that

$$
\begin{equation*}
U=w_{1}\left(Q^{1}\right)+w_{2}\left(Q^{2}\right) \omega^{2} f \tag{1.7}
\end{equation*}
$$

where $Q^{1}$ and $Q^{2}$ are defined by the formulas (1.3) and $w_{1}, w_{2}$ are arbitrary continuous functions.

Proof. The sufficiency of condition (1.7) is obvious (the separate coordinates $u=u\left(Q^{1}\right)$, $v=v\left(Q^{2}\right)$ ). We shall show the necessity. Setting in (1.4) $x=\int \omega d Q^{1}, y=Q^{2}$, we obtain

$$
\begin{equation*}
d s^{2}-\lambda(x)\left(d x^{2}+d y^{2}\right), \quad \lambda-1 /\left(\omega^{2} f\right) \tag{1.8}
\end{equation*}
$$

According to the converse of the Liouville's theorem /11/, if the generalized coordinates in the Hamilton-Jacobi equation of the system are separable, then normal coordinates $u$ and $v$ exist in which

$$
\begin{equation*}
d s^{2}=[\mu(u)+v(v)]\left(d u^{2}+d v^{2}\right), \quad U=\frac{m_{1}(u)+m_{3}(v)}{\mu+v} \tag{1.9}
\end{equation*}
$$

From (1.8) and (1.9) it follows that the transformation $u=u(x, y), v=v(x, y)$ is conformal, therefore we have ( $\varepsilon$ is equal to +1 or -1 )

$$
\begin{equation*}
\frac{\partial u}{\partial x}=\varepsilon \frac{\partial v}{\partial y}, \quad \frac{\partial u}{\partial y}=-\varepsilon \frac{\partial v}{\partial x} \tag{1.10}
\end{equation*}
$$

and in addition

$$
\begin{equation*}
\left.(\mu+v) \left\lvert\,\left(\frac{\partial u}{\partial x}\right)^{2}+\left(\frac{\partial v}{\partial x}\right)^{2}\right.\right]=\lambda \tag{1.11}
\end{equation*}
$$

The geodesic equations in $V_{2}$ with the Liouville's metric (1.9), admit the integral
According to (1.6) we have $\left(J=\lambda y^{\circ}\right)$ ( $\left.\mu+v\right)\left(v u^{\cdot 2}-\mu v^{\circ 2}\right)=$ const

$$
\begin{align*}
& (\mu+v)\left[v\left(\frac{\partial u}{\partial x}\right)^{2}-\mu\left(\frac{\partial v}{\partial x}\right)^{2}\right]=c_{1} \lambda,(\mu+v)\left[v\left(\frac{\partial u}{\partial y}\right)^{2}-\mu\left(\frac{\partial v}{\partial y}\right)^{2}\right]=  \tag{1.12}\\
& c_{1} \lambda+c_{2} \lambda^{2}, \quad(\mu+v)\left(v \frac{\partial u}{\partial x} \frac{\partial u}{\partial y}-\mu \frac{\partial v}{\partial x} \frac{\partial v}{\partial y}\right)=0
\end{align*}
$$

Clearly, the equations (1.10)-(1.12) are compatible only when $u=\varepsilon b x+l, v=b y+r$ or $u=$ $-\varepsilon b y+l, \quad v=b x+r(b, l, r$ are arbitrary constants and $b \neq 0)$, and in both cases $\mu+v=$ $\lambda / b^{2}$. But then from the expression (1.9) for the force function $U$ follows (1.7), Q.E.D.

As an example, we shall consider a heavy double pendulum moving in a vertical plane. A weightless rod $O A$ of length $l_{1}$ rotates about a fixed point $O$ and carries a point mass $m_{1}$ attached to its end. A second weightless rod $A B$ of length $l_{2}$ rotates about the point $A$, with a point mass $m_{2}$ attached to its free end. Let us denote by $\alpha$ the angle between the rod $O A$ and the vertical, and by $\beta$ the angle between the $\operatorname{rod} A B$ and the continuation of the segment $O A$ past the point $A$. The double kinetic energy of the system is

$$
2 T=\left[m_{1} l_{1}{ }^{2}+m_{2}\left(l_{1}{ }^{2}+2 l_{1} l_{2} \cos \beta+l_{2}{ }^{2}\right)\right] \alpha^{\prime 2}+2\left(m_{2} l_{1} l_{2} \cos \beta+m_{2} l_{2}{ }^{2}\right) \alpha^{\prime} \beta^{\prime}+m_{2} l_{2}{ }^{2} \beta^{\prime 2}
$$

and the force function is

$$
\begin{equation*}
U=\left(m_{1}+m_{2}\right) g l_{1} \cos \alpha+m_{2} g l_{2} \cos (\alpha+\beta) \tag{1.13}
\end{equation*}
$$

The Riemannian manifold $V_{2}$ of the system represents a torus $\{\alpha, \beta \bmod 2 \pi\}$ with the kinematic linear element

$$
\begin{align*}
& d s^{2}=E \gamma^{2}+\left(m_{2} l_{2}{ }^{2}-G^{2} / E\right) \beta^{\prime 2}  \tag{1.14}\\
& E=m_{1} l_{1}{ }^{2}+m_{2}\left(l_{1}^{2}+2 l_{1} l_{2} \cos \beta+l_{2}^{2}\right), G=m_{2} l_{1} l_{2} \cos \beta+m_{2} l_{2}{ }^{2} \\
& \gamma=\alpha+x, \quad x=\int \frac{G}{E} d \beta
\end{align*}
$$

We have

$$
\begin{aligned}
& K=\frac{m_{1} v}{\left(l_{1} l_{2} r\right)^{4}}, \quad v=\cos \beta, \quad r=\left(m_{1}+m_{2}-m_{2} v^{2}\right)^{1 / 2}, \quad z=m_{1}+m_{2}+3 m_{2} v^{2} \\
& \Delta_{1} K=\frac{\left(m_{1} z\right)^{2} E\left(1-v^{2}\right)}{m_{2}{ }^{2}\left(l_{1} l_{2}\right)^{12} r^{18}}, \quad \Delta_{2} K=\frac{m_{1} \sqrt{1-v^{2}}}{m_{2}^{2}\left(l_{1} l_{2}\right)^{8} r} \frac{\partial}{\partial v}\left(\frac{z E \sqrt{1-v^{2}}}{r^{9}}\right)
\end{aligned}
$$

The left-hand part of the second relation of (1.5) (say) is equal, when $\beta=0$, to

$$
18 \frac{\left(m_{1}+4 m_{2}\right)^{3}\left[m_{1} l_{1}^{3}+m_{2}\left(l_{1}+l_{3}\right)^{2}\right]^{3}}{m_{1}^{18} m_{2}{ }^{6}\left(l_{1} l_{2}\right)^{16}}>0
$$

Consequently the differential equations of the geodesic lines in $V_{2}$ admit exactly two independent quadratic integrals, $T=$ const and $\left(\partial T / \partial x^{\prime}\right)^{2}=$ const. In accordance with the theorem proved above, the necessary and sufficient condition for the scparable gencralized coordinates to exist is, that the force function has the form (see (1.4) and (1.14))

$$
\begin{equation*}
\left(m_{1}+m_{2}\right) g l_{1} \cos \alpha+m_{2} g l_{2} \cos (\alpha+\beta)=f_{1}(\beta)+f_{2}(\gamma) E^{-1}(\beta) \tag{1.15}
\end{equation*}
$$

However the identity (1.15) is impossible when the parameters of the system have nonzero values. Indeed, differentiating (1.15) with respect to a we obtain $f_{1}+\left(f_{2}+f_{2}{ }^{\prime \prime}\right) E^{-1}=0$, and this implies

$$
\left(m_{1}+m_{2}\right) g l_{1} \cos \alpha+m_{2} g l_{2} \cos (\alpha+\beta)=E^{-1}\left[C_{1} \cos (\alpha+x)+C_{2} \sin (\alpha-\alpha)\right]
$$

where $C_{1}$ and $C_{2}$ are constants. The coefficients of both parts of this identity should coincide for $\sin \alpha$ (or the corresponding $\cos \alpha$ )

$$
\begin{aligned}
& \left(m_{1} \div m_{2}\right) g l_{1}+m_{2} g l_{2} \cos \beta=E^{-1}\left(C_{1} \cos \alpha+C_{2} \sin x\right) \\
& -m_{2} g l_{2} \sin \beta-E^{-1}\left(\quad C_{1} \sin x+C_{2} \cos x\right)
\end{aligned}
$$

Combining the squeres of the resulting relations term by term, we obtain

$$
\left(m_{1}+m_{2}\right)^{2}\left(g l_{1}\right)^{2}+\left(m_{2} g l_{2}\right)^{2}+2\left(m_{1}-m_{2}\right) m_{2} g^{2} l_{1} l_{2} \cos \beta=\left[m_{1} l_{1}^{2}+m_{2}\left(l_{1}^{2}+2 l_{1} l_{2} \cos \beta+l_{2}^{2}\right)\right]^{-2}\left(C_{1}^{2}+C_{2}^{2}\right)
$$

which is not an identity. Thus the generalized separable coordinates cannot in this case be found and the equations of the system motion admit only one quadratic integral, namely the energy integral.
2. Let us consider a conservative system with any number of degrees of freedom and the Hamiltonian function of the form (0.4). We shall call such system an invertible system if the covector

$$
\left\{a_{i}=\sum_{j=1}^{n} a_{i j} a^{j}\right\}
$$

is of the gradient type, i.e. $\quad \partial_{i} a_{j}-\partial_{j} a_{i}=0(i<j=2, \ldots, n)$. Since the presence of the separable generalized coordinates in the system implies that its Hamiltonian function can be written in the form ( 0.5 ), the following theorem holds.

Theorem 2. Generalized coordinates exist in an irreversible system only when the system has cyclic coordinates (implicit or explicit).

Certain invariant criteria of existence of the implicit cyclic coordinates in conservative systems are known. The corresponding criterion for $n=2$ was obtained in /14-16/. Let us denote by $\Phi$ one of the functions

$$
\begin{equation*}
K, \operatorname{rot} e=\delta^{-1}\left(\partial_{1} a_{2}-\partial_{2} a_{1}\right), a \tag{2.1}
\end{equation*}
$$

where $K$ is the Gaussian curvature of the manifold with the metric (1.1), and $\delta^{2}=a_{11} a_{22}-\left(a_{12}\right)^{2}$. The necessary and sufficient condition for the cyclic coordinate to exist in the neighbortuod of any regular point of the function $\Phi$ is, that the invariants (2.1) and

$$
\Delta_{1} \Phi=\sum_{i, j=1}^{2} a^{i j} \partial_{i} \Phi \partial_{j} \Phi, \quad \Delta_{2} \Phi=\frac{1}{\delta} \sum_{i, j=1}^{2} \partial_{i}\left(\delta a^{i j} \partial_{j} \Phi\right)
$$

are functions of $\Phi$. The following obvious assertion is also true: if $q^{*}$ is a critical point of any function $\Phi$ chosen from the invariants (2.1) and the solution of the equation $\Phi(q)=$ $\Phi\left(q^{*}\right)$ in an arbitrarily small neighborhood of the point $q^{*}$ is not a manifold, then the system has no cyclic coordinates in the neighborhood of this point. When $n \geqslant 2$, the system has no cyclic coordinates of the differential scalar invariants of the form

$$
\sum_{i, j=1}^{n} a_{i j} d q^{i} d q^{j}
$$

and the vector fields $\left\{a_{i}\right\},\left\{\partial_{j} a\right\}$ contain $n$ functionally independent quantities.
We use the following examples to illustrate the applications of Theorem 2.

Example 1. Restricted, circular three-body problem. The Lagrange's function of the planetoid (a particle with small mass) is written in the form /17/

$$
\begin{aligned}
& L=\frac{1}{2}\left(x^{2}+y^{2}\right)+\omega\left(x y^{-}-y x^{2}\right)+\frac{1}{2}\left(\omega^{2}\left(x^{2}+y^{2}\right)+\frac{\gamma \alpha}{r}+\frac{\gamma \beta}{s}\right. \\
& r^{2}=(x+a)^{2}+y^{2}, \quad s^{2}=(x-b)^{2}+y^{2}, \quad \omega^{2}=\frac{\gamma(\alpha+\beta)}{(\alpha+b)^{3}}
\end{aligned}
$$

where $\alpha, \beta, \gamma, a, b$ are positive constants. As we know, the critical points of the generalized potential energy

$$
\Phi=-\frac{1}{2} \omega^{2}\left(x^{2}+y^{2}\right)-\frac{\gamma \alpha}{r}-\frac{\gamma \beta}{s}
$$

distributed on the straight line $y=0$ are saddle points, therefore it is impossible, at least in the neighborhood of each of these points, to introduce the generalized coordinates one of which would be cyclic. Generally speaking, the system has no locally cyclic coordinates. Indeed, according to $/ 14,16 /$, the necessary condition for a cyclic coordinate to exist is, that the force lines be the geodesics of an Euclidean plane. This means that for every regular point $\left(x_{0}, y_{0}\right)$ of the function $\Phi$ constants $A, B\left(A^{2}+B^{2} \neq 0\right)$ exist such that

$$
A \frac{\partial \Phi}{\partial x}+B-\frac{\partial \Phi}{\partial y}=0
$$

along the straight line $A\left(x-x_{0}\right)+B\left(y-y_{0}\right)=0$. Since

$$
\frac{\partial \Phi}{\partial x}=-\omega^{2} x+\frac{\alpha(x+a)}{r^{3}}+\frac{\beta(x-b)}{s^{3}}, \quad \frac{\partial \Phi}{\partial y}=-\omega^{2} y+\frac{\alpha y}{r^{3}}+\frac{\beta y}{s^{3}}
$$

the condition at infinity implies $A x \mid B y=0$. But then we have

$$
A\left(\frac{\alpha a}{r^{3}}-\frac{\beta b}{s^{3}}\right)=0
$$

which is a contradiction. By Theorem 2, any generalized coordinates of the planetoid cannot therefore be locally separated.

A different proof of this assertion exists /7/ based on the fact that the equation

$$
\left(\frac{\partial V}{\partial x}\right)^{2}+\left(\frac{\partial V}{d y}\right)^{2}=h
$$

can be integrated by separating the variables only in the elliptical and degenerate elliptical coordinates.

Example 2. A heavy rigid body bounded by an arbitrary convex surface slides without friction along a stationary horizontal $O x y$-plane. We introduce the moving axes $G \xi \eta_{\mathrm{g}}$ coinciding with the contral axes of inertia of the body, where $G$ is its center of mass. We direct the $O z$-axis of the fixed $O x y z$ coordinate system vertically upwards. The distance between the point $G$ and the supporting plane will be denoted by $s$. We choose, as the generalized coordinates of the body, the coordinates $x$ and $y$ of $G$, and the Euler angles $\varphi, \psi, \theta$ of rotation of the trihedron $G \xi \eta \zeta$ relative to the $O x y z$ axes. Clearly, $z=z(\varphi, \theta)$ and the coordinates $x, y$ and $\psi$ are cyclic. Ignoring the cyclic coordinates we obtain the Routh function of the reduced system

$$
\begin{aligned}
& R=R_{2}+R_{1}+R_{0} \\
& R_{2}=\frac{1}{2}\left\{\left[\frac{C(F+C) \sin ^{2} \theta}{v}+M\left(\frac{\partial z}{\partial \varphi}\right)^{2}\right] \varphi^{2}-2\left[\frac{C(A-B) \sin 2 \theta \sin 2 \varphi}{4 v}-\right.\right. \\
& \left.\left.M \frac{\partial z}{\partial \varphi} \frac{\hat{\sigma}_{z}}{\partial \theta}\right] \varphi \cdot \theta_{-}^{-}\left[F+C-\frac{(A-B)^{2} \sin ^{2} \theta \sin ^{2} 2 \varphi}{4 v}+M\left(\frac{\partial z}{\partial \theta}\right)^{2}\right] \theta^{*}\right\} \\
& n_{1}-\frac{i}{v}\left[\varphi C \cos \theta+\frac{1}{2} \theta \cdot(A \quad B) \sin \theta \sin 2 \varphi\right] \\
& R_{0}=M g z-\frac{1}{2} \frac{j^{2}}{v}-\frac{M}{2}\left[(x)_{0^{2}}^{2}+\left(y^{2}\right)_{0}{ }^{2}\right] \\
& F=A \sin ^{2} \varphi+B \cos ^{2} \varphi-C, \nu=F \sin ^{2} \theta+C
\end{aligned}
$$

Here $j,\left(x^{*}\right)_{0},\left(y^{\circ}\right)_{0}$ are the integration constants, $A, B$ and $C$ are the principal moments of inextia and $M g$ denotes the weight of the body.

Excluding from our discussion the case $z \equiv$ const (symmetrical sphere), we establish whether separate generalized coordinates exist in the Hamilton-Jacobi equation of the reduced system.

For this the presence of an implicit cyclic coordinate is necessary. Let us assume that such a coordinate exists. Then by virtue of the arbitrariness of the values of the constant $j$ the variables $z$ and 0 must be functions of the position coordinate. This means that $z=z(v), F \not \equiv 0$. On the other hand, the invariant rot must also be a function of the position coordinate , and it is therefore necessary that the Jacobian

$$
\begin{equation*}
\frac{D(\operatorname{rot} e, v)}{D(\varphi, \theta)} \equiv 0 ; \quad \operatorname{rot} e=\frac{1}{\delta}\left(\frac{\partial e_{2}}{\partial \varphi}-\frac{\partial e_{1}}{\partial \theta}\right) \tag{2.2}
\end{equation*}
$$

where $e$ is a covector with components

$$
e_{1}=\frac{C \cos \theta}{v}, \quad e_{2}=\frac{(A-B) \sin \theta \sin 2 \varphi}{2 v}
$$

and $\delta^{2}$ is the determinant of the form $2 R_{2}$ quadratic in' 9 and $\theta$. Carrying out the calculations we find, that the left-hand part of (2.2) is

$$
\begin{aligned}
& (A-B)\left[p_{6}(w, v, A-B, C)+M(d z / d v)^{2} P_{8}(u, v, A-D, C)\right]\left(x^{3} F^{2} V_{v)^{-1}}\right. \\
& w=\sin ^{2} \varphi, \quad x=\delta / \sin \theta
\end{aligned}
$$

where $P_{\mathrm{B}}$ and $P_{g}$ are certain sixth and eighth order polynomials in $w$, with the coefficients defined for all possible values of the parameters and $v \neq 0$. The coefficient accompanying the highest power of $w$ in $p_{B}$ is $8(A-B)^{8} y^{2}$.

Thus we see that the separable coordinates can exist in the reduced system only when $A=B$. But when $A=B$, we have $z=f(\theta)$ (i.e. the body is bounded by a surface of revolution) and the vatiables $t, \psi$ and, $\theta$ are separable.

## REFERENCES

1. LEVI-CIVITA T., Sulla integrazione della equazione di Hamilton-Jacobi per separazione di variabili. Math. Ann. 1904. B.59, s.383-397; Opere matematiche, t.2, (1901-1907), Bologna, Zanichelli, 1956.
2. FORBAT N.H., Sur la separation des variables dans l'equation de Hamilton-Jacobi d'un système non conservatif. Bull. Cl.sci.Acad.roy.Belgique, $5^{e}$ serie,t. $31,1944$.
3. PRANGE G., Die allgemeinen Integrationsmethoden der analytischen Mechanik. Encyklopädie Math. Wiss. B. IV, 12u, 13, Teil lmAbt. 2, H.4, Leipzig, Teubner, 1935.
4. HUAUX A., Sur la séparation des variables dans l'equation aux dérivées partielles de Hamilton-Jacobi. Ann. Mat. pura ed appl. t. 108, 1976.
5. JACOBI C.G.J., Vorlesungen uber Dynamik nebst fünf hinterlassenen Abhandlungen. Berlin, G. Reimer, 1866.
6. EISENHART L.P., Separable systems of Stäskel. Ann. Math. Vol. 35, No. 2, 1934.
7. IAROV-IAROVOI M.S., On integrating the equations of motion of a material point using the method of variable separation. V sb: Tr. mezhvuz. konf. po prikladnoi teorii ustoichivosti dvizheniia i analit. mekhanike. Kazan', 1962. Izd-vo Kazansk. aviats. in-ta, 1964.
8. IAROV-TAROVOI M.S., Integration of the Hamilton-Jacobi equation by the method of separation of variables. PMM Vol.27, No.6, 1963.
9. POINCARÉ H., Les méthodes nouvelles de la Mécanique Céleste. t. 1. Paris, Gauthier-villars, 1892.
10. KOZLOV V.V., Quantitative Analysis Methods in the Solid Body Dynamics. Moscow, Izd-vo MGU, 1980.
11. BENENTI S.., Proprieta' intrinseche dei sistemi dinamici separabili e condizioni necessarie per l'integrabilita' dell'equazione di Hamilton-Jacobi mediante separazione delle variabíli. In: $3^{\circ}$ Congr. naz. Assoc. ital. mecc. teor. ed appl. Cagliari, 1976., Sez. 1 Bologna 1976.
12. SHULIKOVSKII V.I., Ciassical Differential Geometry in Tensor Formulation. Moscow, Fizmatgiz, 1963.
13. DARBOUX G., Leçons sur la théorie generale des surfaces et les applications géometriques de calcul infinitesimal. Pt.3, Paris, Gauthier-Villars, 1894.
14. RICCI G.and LEVI-CIVITA T., Methodes de calcul différentiel absolu et leurs applications. Math. Ann. 1901, B. 54.s. 128-201; LEVI-CIVITA T. Opere matematiche, T. 1, 1893-1900. Bologna:Zanichelli, 1954.
15. SYNGE J.L., On the geometry of dynamics. Philos. Trans. Roy. Soc. London, Ser.A, Vol. 226, No. 637, 1926.
16. SUMBATOV A.S., On the cyclic coordinates of conservative dynamic systems with two degrees of freedom. V sb.:Teoriia ustoichivosti i ee prilozheniia. Novosibirsk, NAUKA, 1979.
17. PARS L.A., A treatise on analytical dynamics. London, Heinemann, 1965.

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